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#### Other Recent Work on HPC

Anomalous transport in magnetically confined plasmas is believed to be driven by microinstabilities, which are fueled by the free energy provided by density and temperature gradients [1] . Some well known instabilities include ion temperature gradient (ITG) modes, trapped electron modes (TEM), electron temperature gradient (ETG) modes, etc. which are driven by tempera-ture gradient. However, there are modes that are driven purely by density gradient, one such example is universal drift mode (UDM). Similar to UDM, another mode driven by the density gradient present in trapped electrons is Ubiquitous mode (UM). UM rotates in ion diamagnetic drift direction in contrast to electron modes such as ETG, TEM, UDM, etc. Generally, long wavelength modes alone do not adequately account for all energy losses in anomalous transport. Consequently, the contributions from short to intermediate wavelength modes, such as Ubiquitous Modes (UMs), cannot be disregarded. Recent observations, particularly the flat temperature profiles alongside finite density gradients reported in LTX tokamak experiments [2,3], underscore the potential role of these modes in driving plasma instability. Furthermore, recent studies employing the ballooning approximation for tokamak plasmas have indicated that the onset of UMs under specific operating parameters can lead to a significant increase in plasma diffusion [4]. In present work, we study the linear properties of the UM and its dependencies

on various parameters using a global gyro-kinetic model, GLOGYSTO [5] based on spectral method.

 First ever gyrokinetic study of density gradient driven, trapped particle mode -- Ubiquitous Mode -- in regime relevant for Lithium Tokamak experiments LTX, suggest strong evidence for transport. the plots, the mode is quite global and occupies a substantial fraction of minor radius. These aposteriori sug-

gests that, a global calculation is necessary such as the present one to correctly account for global effects. Mode structure for the conventional TEM is shown in Fig.5 for the mode number, n = 16. A close-up view of the same is shown in Fig.6.

**Dependence on density gradient :** Since the mode is essentially driven by density gradient, it is critical to study the mode's dependence as a function of gradient in density. Figs. 7 & 8 shown the real frequency and growth rate plots as a function of  $\mathbf{R}/\mathbf{L}_n$  respectively. It can be observed from that with



Fig. 5 Global Mode Structure corresponding to n = 16



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0.5

0.7

0.6

0.5

0.3

0.8

0.6

0.4

0.2

-0.2

z/a

C<sub>s</sub>/a

#### Ubiquitous Mode in Tokamaks

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**Fig.1** Real frequency normalised by  $c_s/a$ ,  $c_s$  being the sound speed.

Growth rates and real frequency calculated for different toroidal mode numbers are shown in Figs. 1 & 2. Frequenber  $c_s=\sqrt{k_BT_e/m_i}$  cies are normalised by  $c_s/a$ , where  $c_s=\sqrt{k_BT_e/m_i}$  is the sound speed.

Wave numbers,  $k_{\theta}$  is normalised with  $\rho_{s} = c_{s}/\omega_{ci}$  where  $\omega_{ci}$  =  $eB/m_{i}$  is the ion cyclotron frequency. For low  $k_{0}\rho_{s}(\leq 1.36)$ , real frequency is in electron drift direction, this is the regime of density gradient driven conventional TEM. However, for  $k_{0}\rho_{s} \geq 1.36$ , the mode direction reverses and rotates in the direction of ion diamagnetic drift. This branch of the dispersion curve corresponds to the ubiquitous mode.

The mode structure for ubiquitous mode for mode number,  $\mathbf{n}$  = 30 is shown Fig. 3 and a zoomed in for the same is shown in Fig 4. As can be seen from



 $k_{\theta} \rho_{s}_{1.5}$ 

2.0

Ubiquitous

35

2.5

1.0

TEM

20 25

**Fig.2** Growth rate normalised by  $c_s/a$ ,  $c_s$  being

the sound speed.



Fig. 4 Close-up view of the mode structure



Fig. 6 Close-up view of the mode structure

Profiles and parameters considered for the study tare:











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increase in  $R/L_n$  i.e., increase in density gradient, growth increases almost linearly. Real frequency scales with diamagnetic drift frequency which is proportional to density gradient as the temperature profile is considered flat, therefore a linear variation in real frequency is observed with density gradient. For the results shown in the article, in total, 12-15 node-hours of ANTYA were used.

**Dependence on**  $\eta = L_n/L_T$ : Ubiquitous mode, so far, is studied for flat temperature profile as it is driven solely by density gradient and does not need gradients in the temperature profile. However, in practical scenarios, tokamaks generally exhibit finite temperature gradients. Therefore, the dependence of the mode on temperature gradient is investigated. For this,  $\eta = L_n/L_T$  is varied by keeping  $L_n$  constant; increase in  $\eta$  means increase in temperature gradient.  $L_n$ ,  $L_T$  are kept the same for both of the species, electrons and ions, Real frequency and growth rates are shown in Fig. 9 & 10 respectively. Real frequency is observed to vary linearly with  $\eta$  as expected since it scales with diamagnetic drift. Growth rate, however, shows weak dependence on temperature gradient and the mode is likely to be present in finite temperature gradient scenarios.



Fig. 7 Real frequency with respect to  $R/L_n$ 

In this study, we systematically investigate the linear properties of the ubiquitous mode utilizing a global gyrokinetic model. We examine the mode's dispersion relation and its structural characteristics. The mode structures are found to occupy a substantial portion of the minor radius, underscoring the importance of global calculations. Additionally, analyze the driving mechanisms of the mode by exploring its dependence on density and temperature gradients. Our findings indicate that the primary driving factor for the mode is the density gradient, while the temperature gradient exhibits a weak influence on the growth rate.

#### References:

1. Horton W., Rev. Mod. Phys. 71 735 (1999) 2. Majeski et al Phys. Plasmas 24, 056110

- (2017)
- 3. Boyle et al Phys. Rev. Lett. 119, 015001 (2017)
- 4. Shen et al Nuclear Fusion 59 106011 (2019)
- 5. Brunner Ph. D. Thesis 1701, EPFL, Switzerland (1997)





Fig. 9 Real frequency with respect to  $\eta$ 



Fig. 10 Growth rate with respect to  $\eta$ 

### **HPC Picture of the Month**



#### Pic Credit: Vinod Saini

<u>Title</u>: Iso-contour of the electron phase-space distribution f(x,v,t) for chirp driven system.

**Description:** The parameters for the simulation are as following:- the initial Maxwellian plasma is driven with the drive amplitude  $E_0 = 0.025$  and frequency is swept from  $\omega_d = 1$  to  $\omega_d = 0.5$  for chirp time interval  $\Delta t_d = 200$  with chirp coefficients are  $\zeta = -2.5 \times 10 - 3$ ,  $\eta = 1$ . Late time contour plot of f(x,v,t) at t=2000, where an external downward chirp is applied from the start for a period  $\Delta t_d = 200$ , after the drive is turned off, using PIC solver. Because of inherent noise in particle simulations, small scale structures in electron phase space are highly noisy. These structures are seen at the less populated tail of the distribution function, which contributes to statistical noise.

Solver User: An in house developed 1D3V PIC-MCC solver is used here

Simulation Details: This simulation result is produced using 512 spatial grid points and 4000 velocity grid point. PIC solver is fully ported on GPU and single GPU card on Antya is used for simulation

Visualization: A licensed Matlab is used to plot the simulation data

# Introduction to NVIDIA HPC SDK for High Performance Computing

High-Performance Computing (HPC) has transformed the way to solve scientific, engineering, and industrial problems by using powerful, efficient, and scalable computers that handle large data and perform complex calculations rapidly. The NVIDIA HPC SDK (Software Development Kit) provides developers with a robust platform to build, optimize, and deploy HPC applications. It includes compilers, libraries, and tools that are essential for developing applications that run on NVIDIA GPUs and CPUs.

NVIDIA HPC SDK (<u>Refer here</u>) also supports GPU acceleration using CUDA and other parallel computing standards like OpenACC and OpenMP. This helps developers use the powerful parallel processing abilities of NVIDIA GPUs, greatly improving their applications' performance and efficiency. The SDK works well with popular development environments and supports languages like Fortran, C, and C++.

Additionally, the NVIDIA HPC SDK includes a collection of highly optimized libraries for various mathematical and scientific computations. These libraries, such as cuBLAS for linear algebra, cuFFT for fast Fourier transforms, and cuSPARSE for sparse matrix operations, are designed to deliver higher performance and ease of use. It also supports communication libraries such as OpenMPI, NCCL etc. Below is the list of compilers and libraries available with SDK.

Programming Models	Compilers	Core Libraries	Math Libraries
<ul> <li>Standard C++, Fortran</li> <li>OpenACC &amp; OpenMP</li> <li>CUDA</li> </ul>	<ul><li>NVCC &amp; NVC</li><li>NVC++</li><li>NVFORTRAN</li></ul>	<ul><li>Libcu++</li><li>Thrust</li><li>CUB</li></ul>	<ul> <li>cuBLAS &amp; cuTENSOR</li> <li>cuSPARSE &amp; cuSolver</li> <li>cuFFT &amp; cuRAND</li> </ul>

Source: https://developer.nvidia.com/hpc-sdk

There are two ways to install the NVIDIA HPC SDK: 1) Using Target Environment-based installation, and 2) Container-based installation. To install the NVIDIA HPC SDK in your environment, visit the <u>SDK download page</u>. Select a compatible version of the SDK according to the drivers and CUDA version available with the system. Follow the installation instructions for the target platform and refer to the steps below for installing the SDK into users' environment after downloading and extracting the compatible version from the SDK download page. After installation, detailed instructions regarding environment setup will be shown to the user.



To verify the installation of the SDK toolkit, various example codes related to different libraries and compilers are included with the HPC SDK installation. Users can load the SDK toolkit into their environment using the command mentioned above and run any of the example scripts to verify the installation of the HPC SDK.

#To load the NVHPC SDK module in user environment [user@login1 ~]\$ module load /scratch/scratch\_run/user/NVHPC\_20.7/install/modulefiles/nvhpc/20.7

[user@login1 ~]\$ which nvc

/scratch/scratch\_run/user/NVHPC\_20.7/install/Linux\_x86\_64/20.7/compilers/bin/nvc

#List of example scripts present in example directory

[user@login1 ~]\$ cd /scratch/scratch\_run/user/NVHPC\_20.7/install/Linux\_x86\_64/20.7/examples AutoPar CUDA-Fortran CUDA-Libraries F2003 MPI OpenACC OpenMP README

#Submit an interactive job in any one of the gpu and test the example scripts by traversing into the desired example directory.

[user@gn01 AutoPar]\$ make all [user@gn01 cuBLAS]\$ make all

#### Users may use interactive mode only for testing above installation.

ΝΟΤΕ

The NVIDIA HPC SDK comes in various versions, each supporting a specific set of GPU devices, driver versions, and compute capabilities. Installing an incompatible version may lead to errors related to driver versions, compute capabilities, libraries and more. Therefore, users should refer to the documentation and download the SDK version that is compatible with their GPU device.

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**ANTYA HPC USERS'** 

# **ANTYA Utilization: MAY 2024**

### ANTYA Daily Observed Workload



# **Other Recent Work on HPC**

Excitation of 2-D and 3-D Pinned Solitons in a Flowing Dusty Plasma	Prasanta Amat	MAY 2024
Spontaneous Convective Patterns in a Dusty Plasma	Ankit Dhaka	Total Successful Jobs~ 1522
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Convective cell to shear flow instability in 2D Yukawa liquids driven by Reynolds stress: A first principles study	Rajaraman Ganesh	CPU Cores Amit Singh     GPU Cards Shishir Biswas
Short-pulse laser-cluster interaction in an ambient magnetic field	Kalyani Swain	
Laser-cluster interaction in an ambient magnetic field with circular polarization	Kalyani Swain	Walltime Shishir Biswas
Design Development of Toroidal Field Coil for Prototype Center Stack	Aditya Kumar Verma	Jobs Prince Kumar

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HPC Environment for New Users Available Please send your request to hpcteam@ipr.res.in.hpcusers@ipr.res.in If you wish to contribute an article in GANANAM, please write to us.HFEmail: hpcteam@GANANAM, please write to us.Email: hpcteam@
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